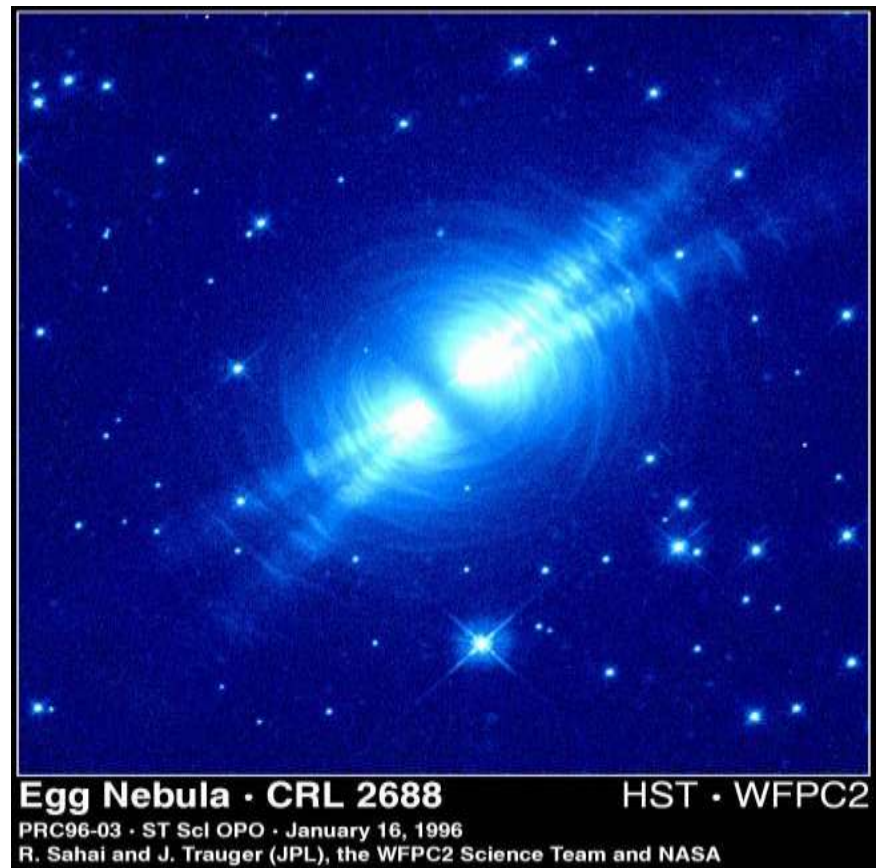
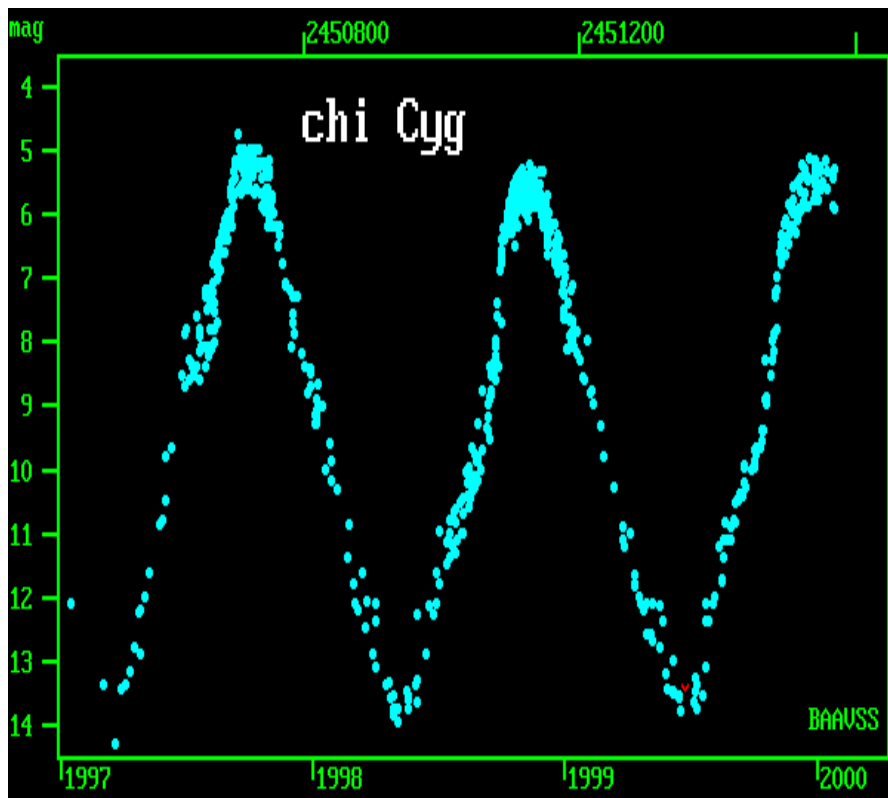
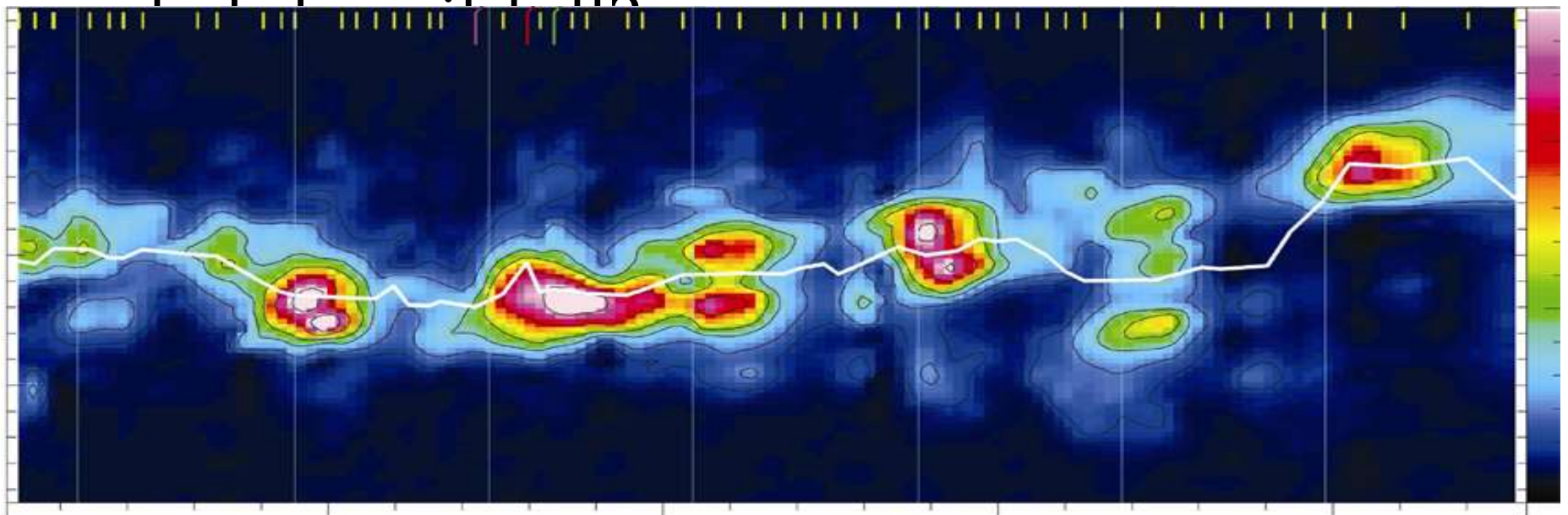
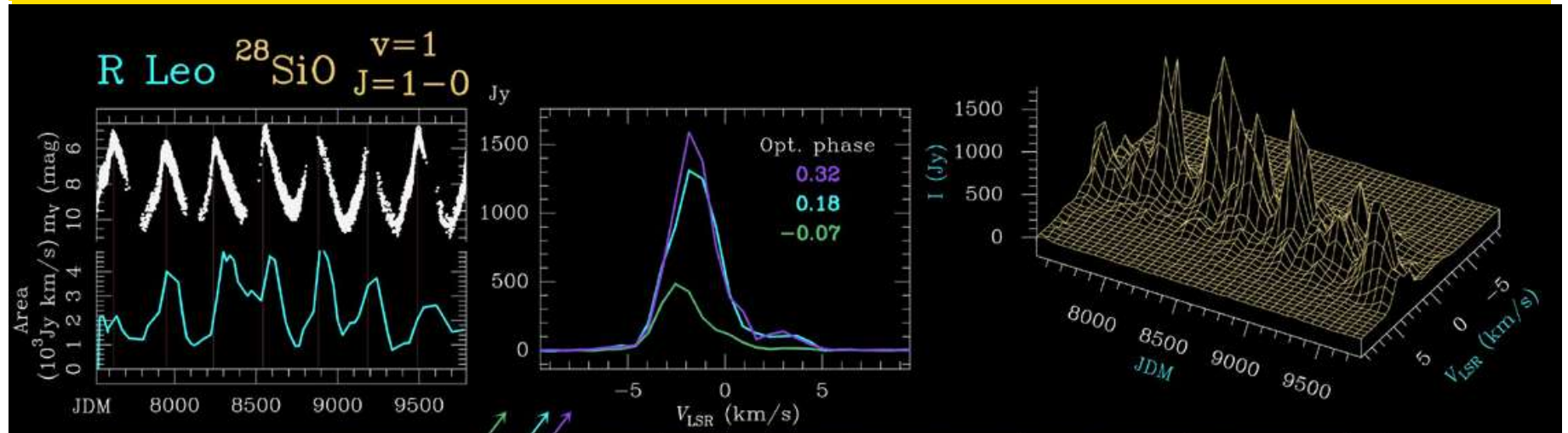


Group C – J.R. Pardo
SiO masers in AGB stars

-
Comparing CRL 2688, CRL 618
and NGC 7027

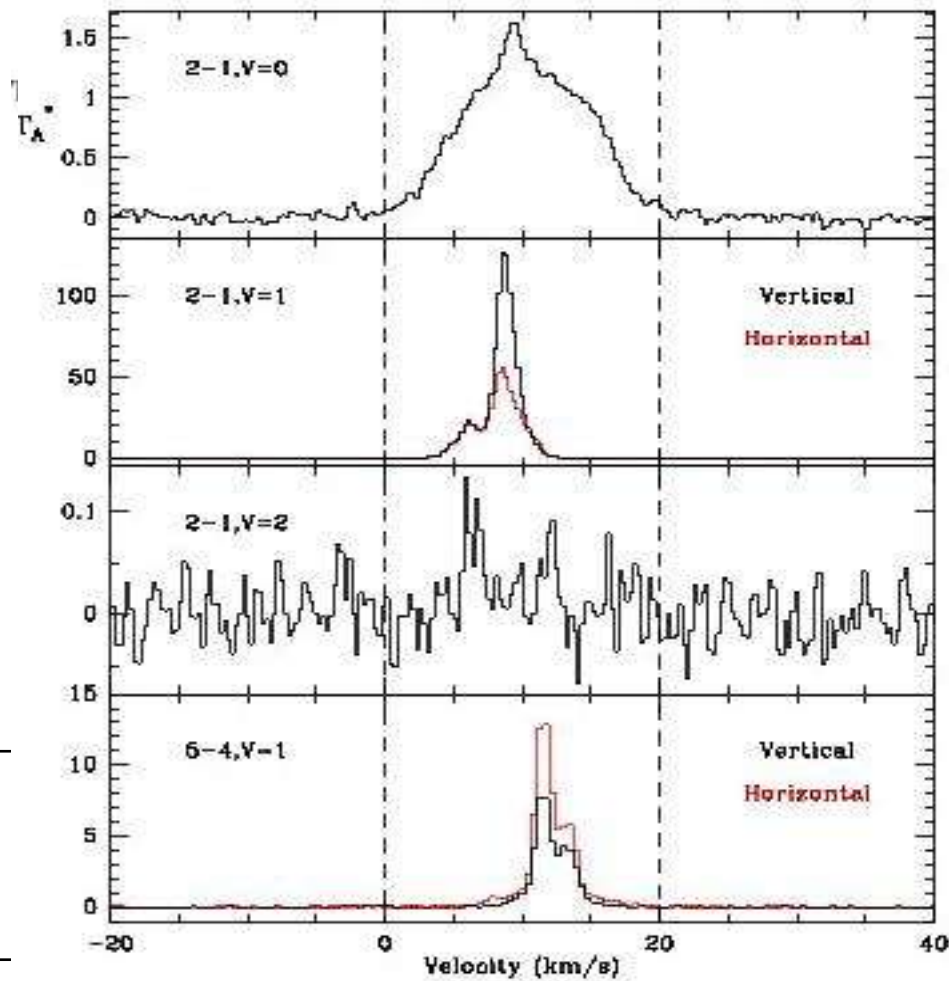


SiO masers

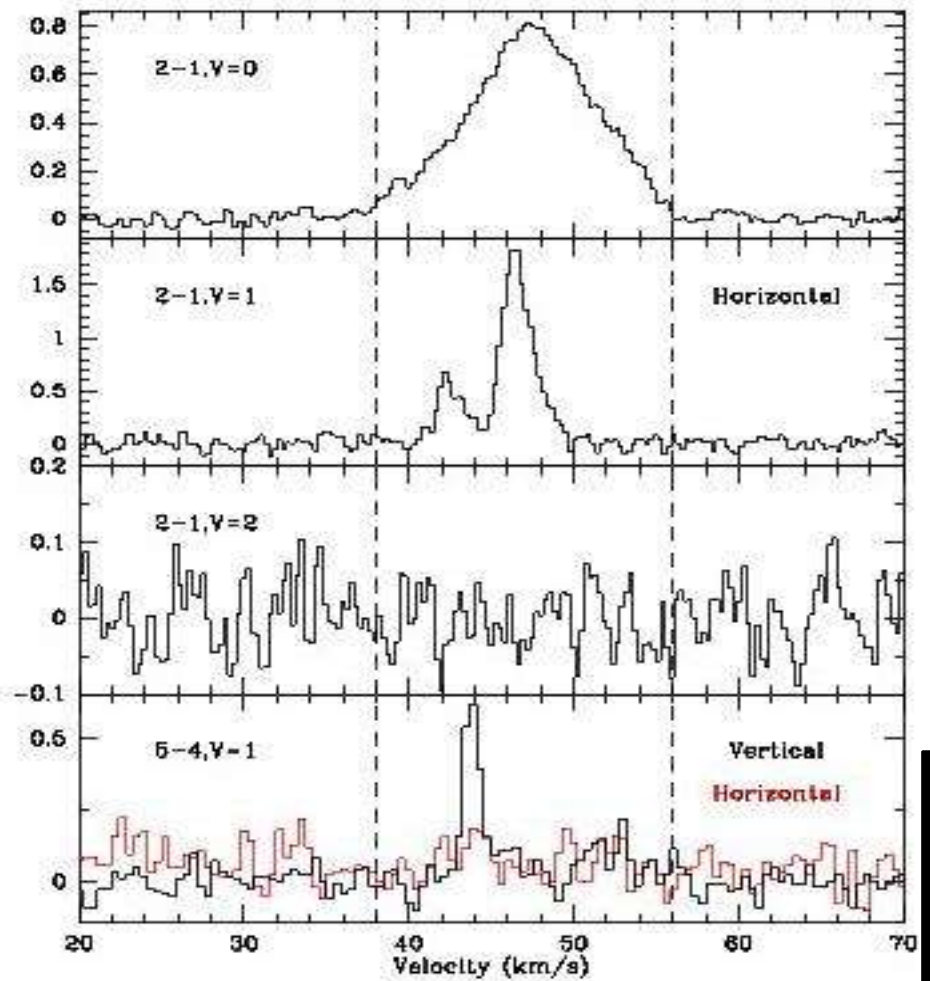


SiO masers

χ Cyg

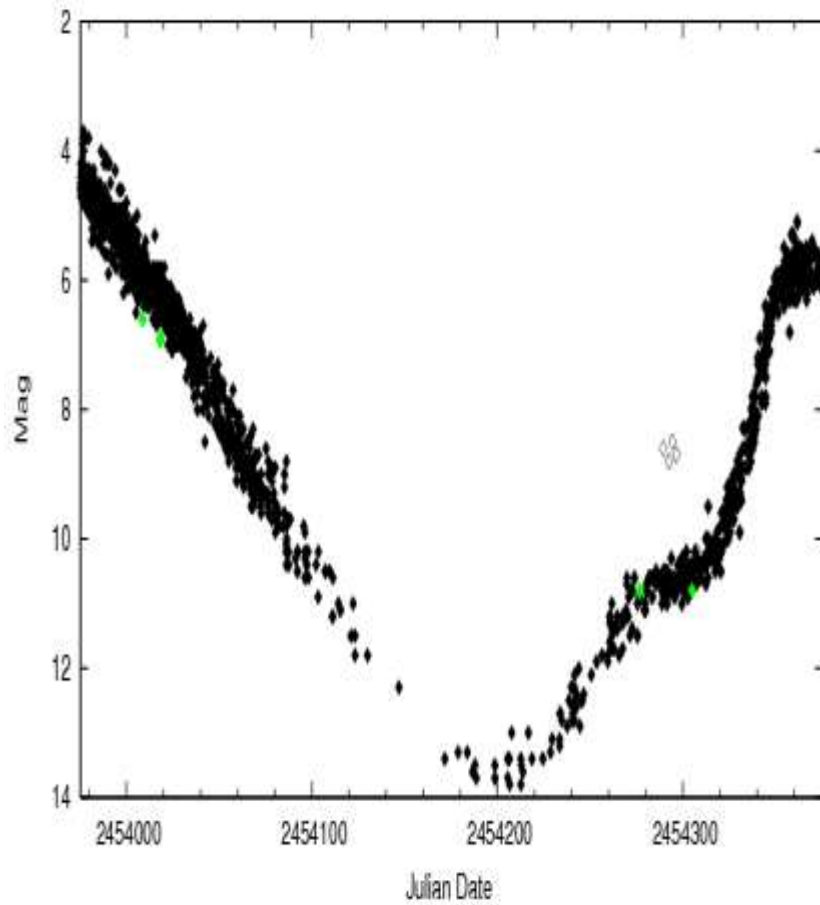


R Aql



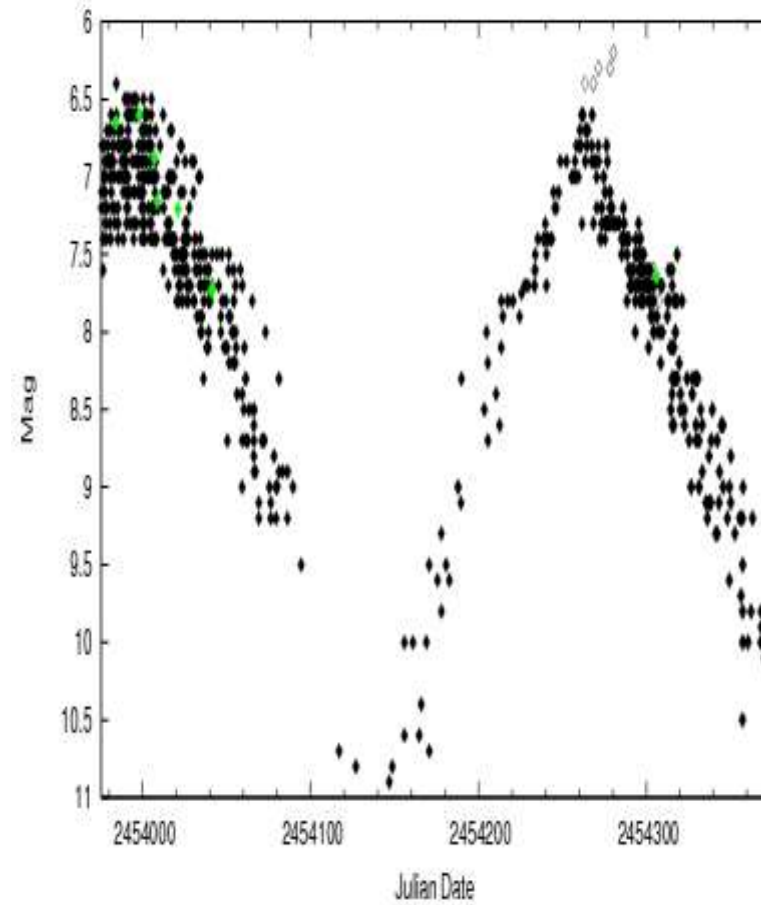
SiO masers

AAVSO DATA FOR CHI CYG - WWW.AAVSO.ORG



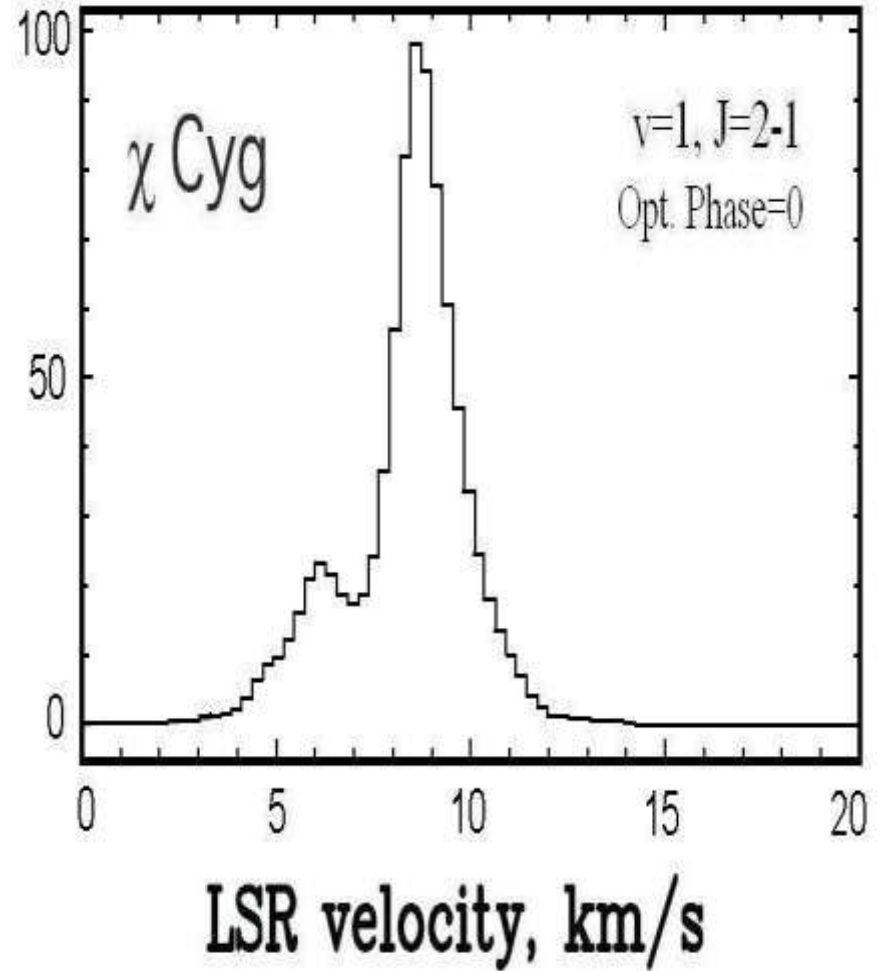
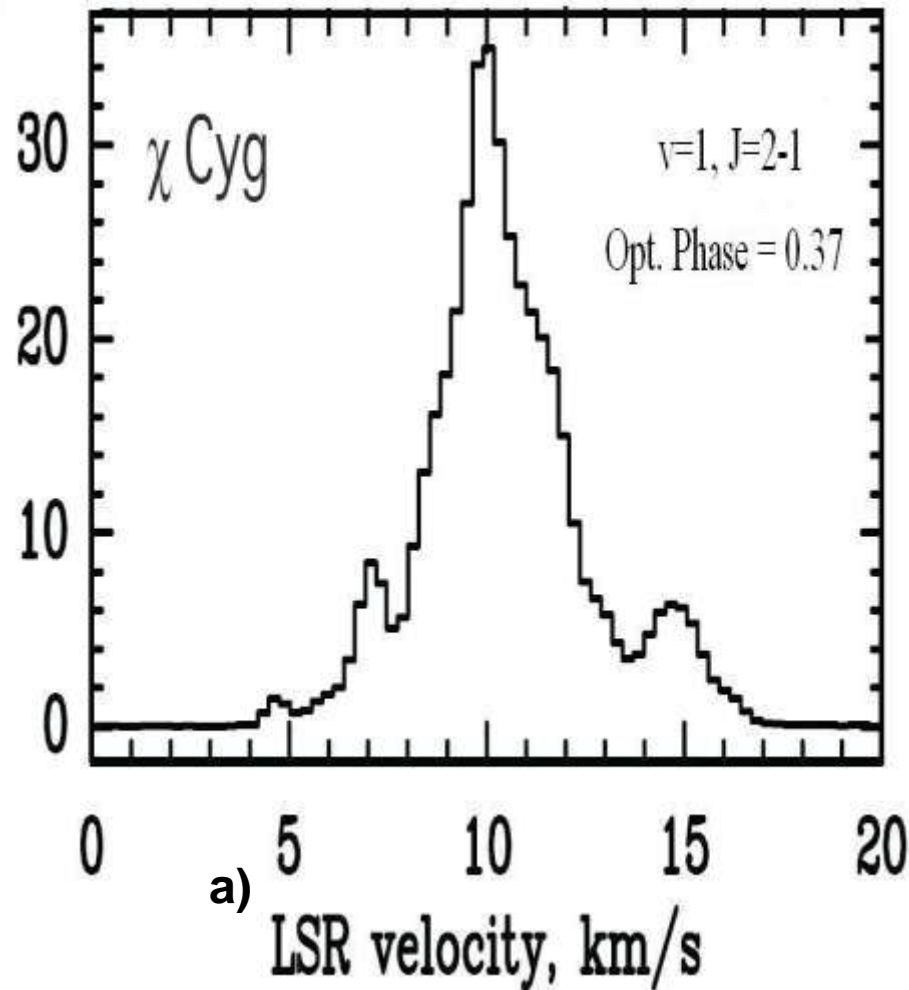
Visual Unvalidated \diamond V \diamond
Visual Validated \blacklozenge

AAVSO DATA FOR R AQL - WWW.AAVSO.ORG



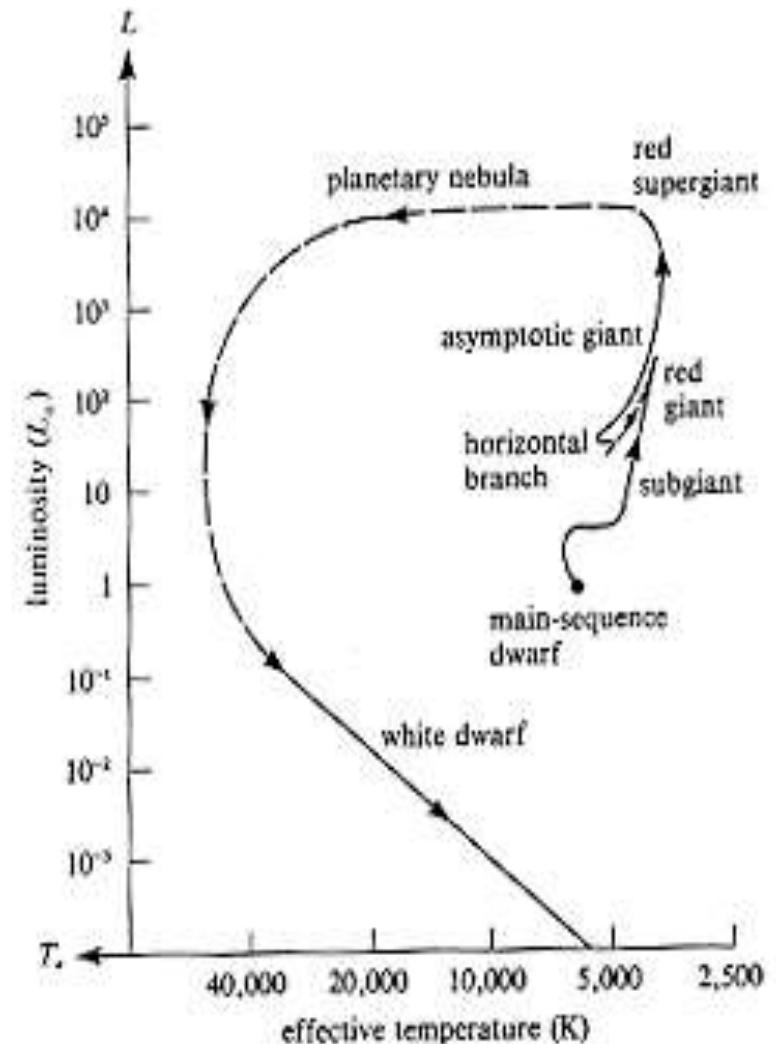
Visual Unvalidated \diamond V \diamond
Visual Validated \blacklozenge

SiO masers



Protoplanetary Nebulae (PPNe)

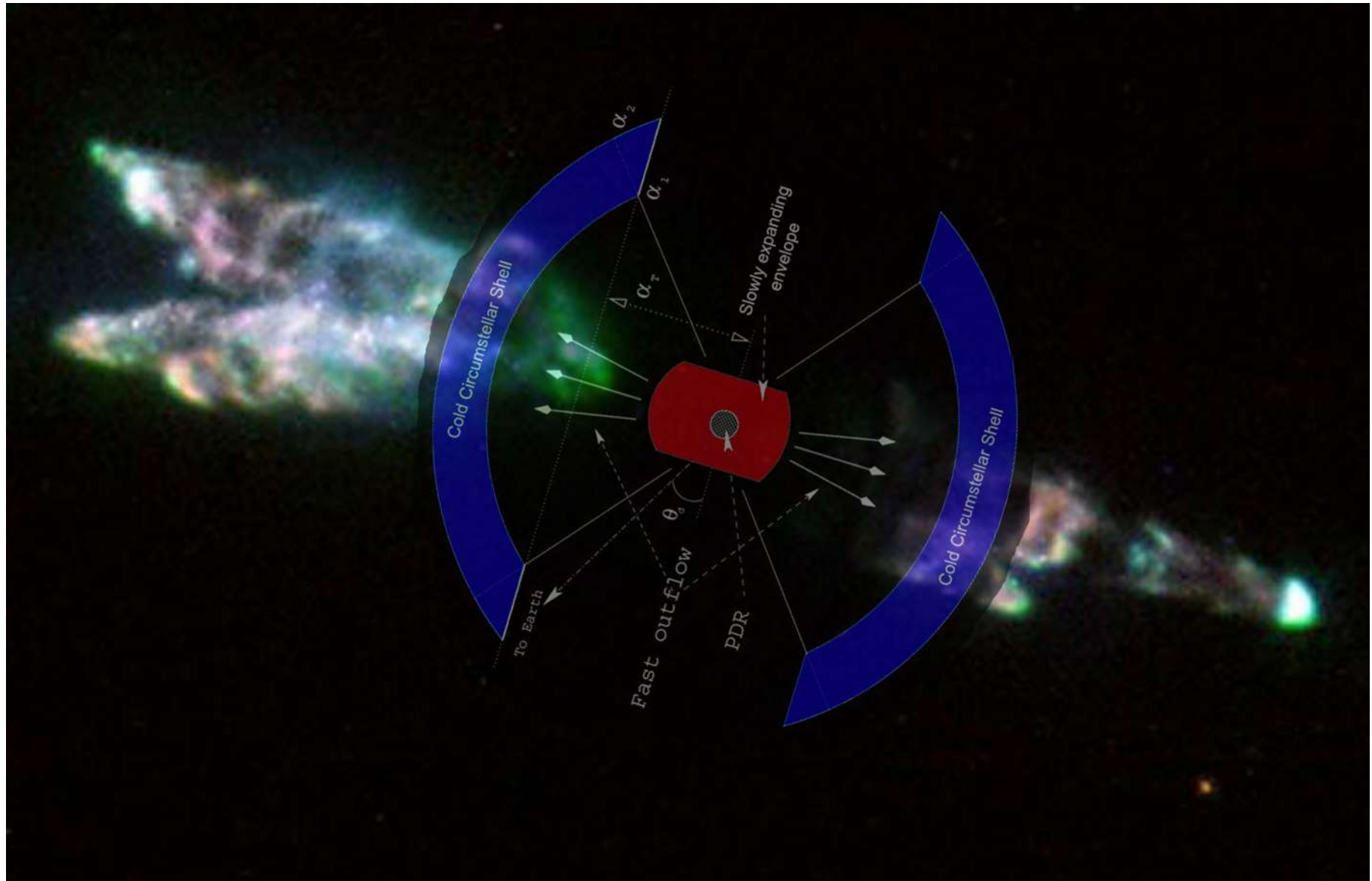
- transition step from Post-AGB to Planetary Nebula and finally White Dwarf ± 1000 yrs
- central star
 - few 10^4 K
 - enshrouded by dust
- complex structure with ejections and jets



HC₃N in CRL 618

- C-rich
- HC₃N
 - linear molecule: equally spaced rotational transitions
2B=9GHz
 - 2 lines in 1mm-window
2 lines in 3 mm-window
 - good tracer for SEE because of many low-lying vibrational states, which can be populated in LTE regime

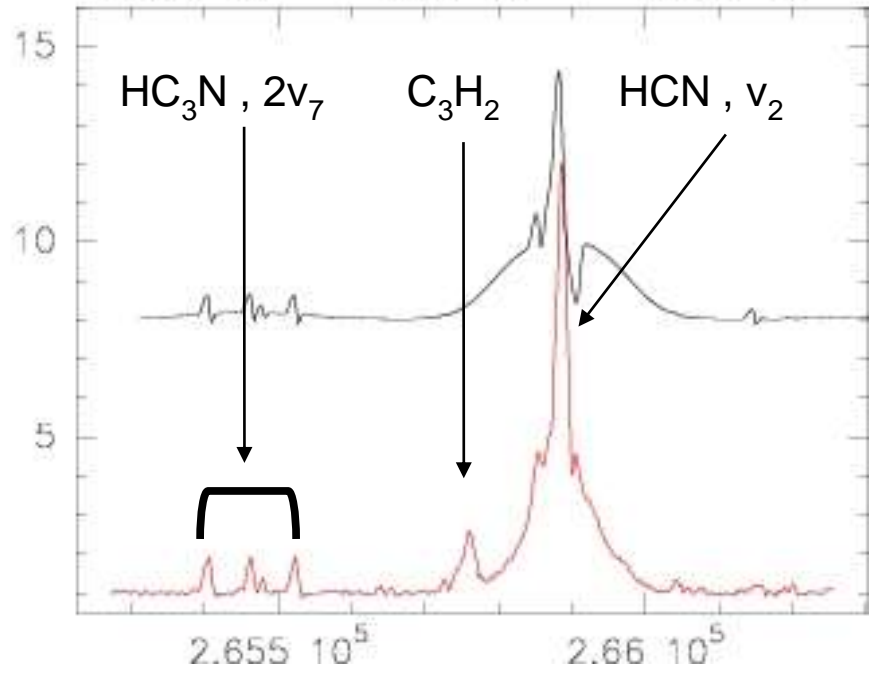
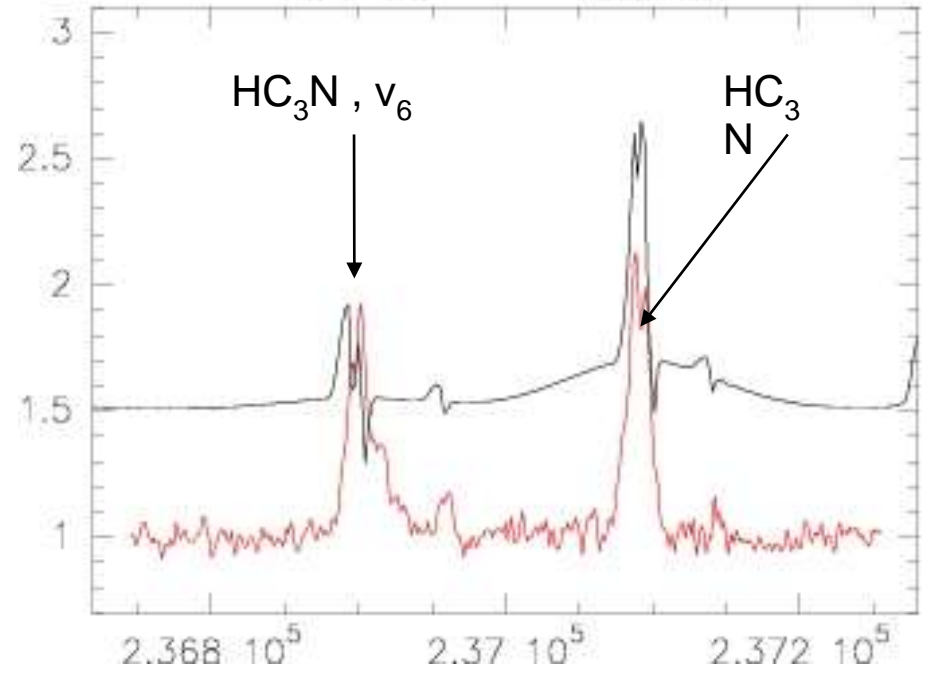
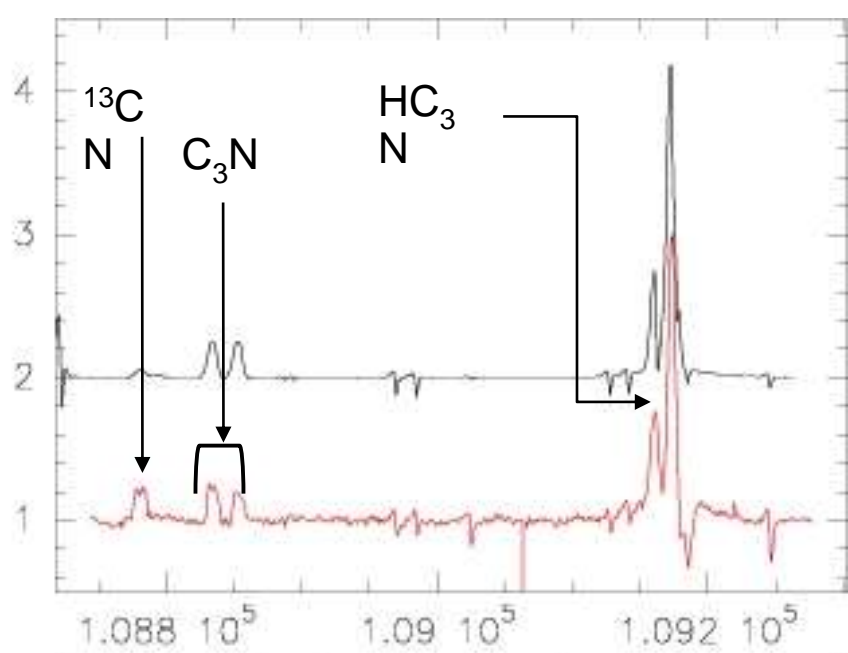
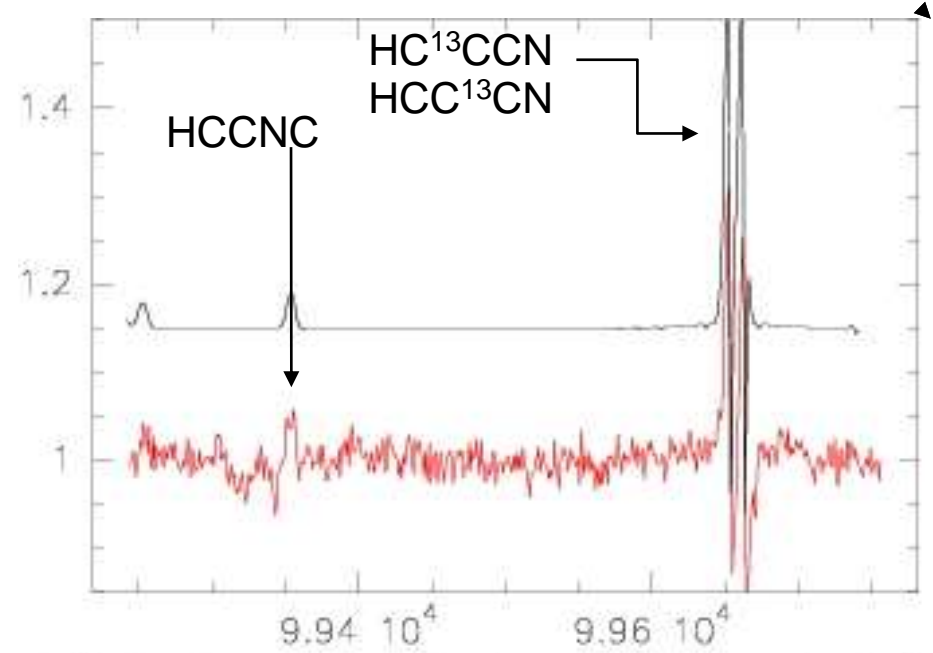
CRL 618: Optical image + structure of molecular outflows



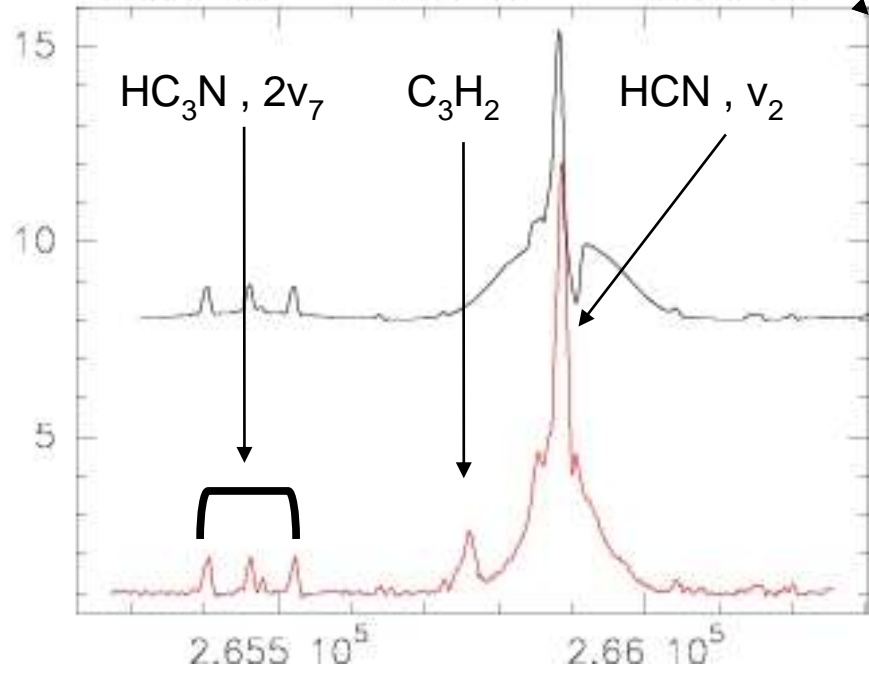
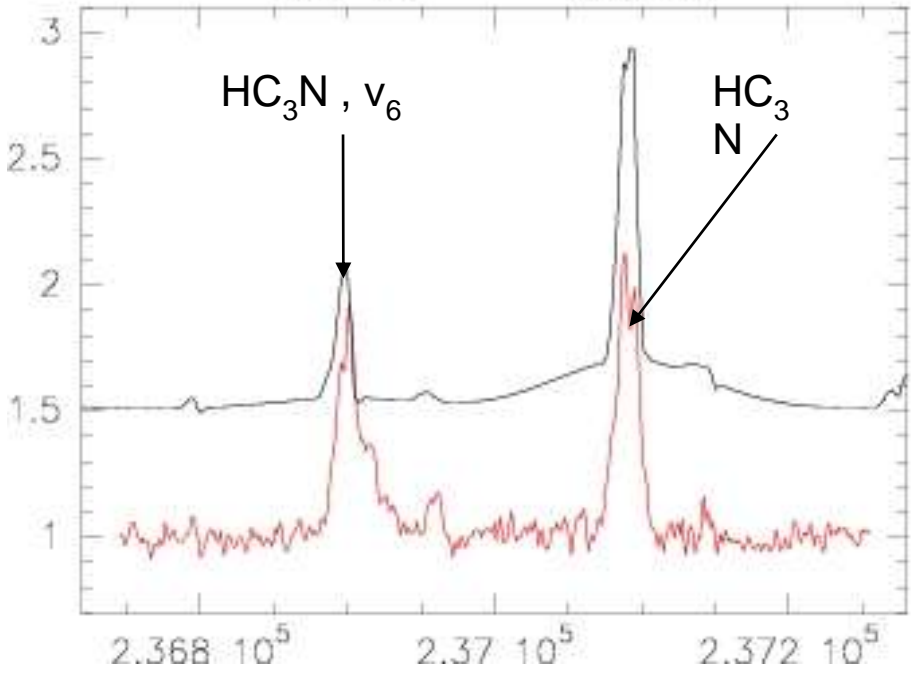
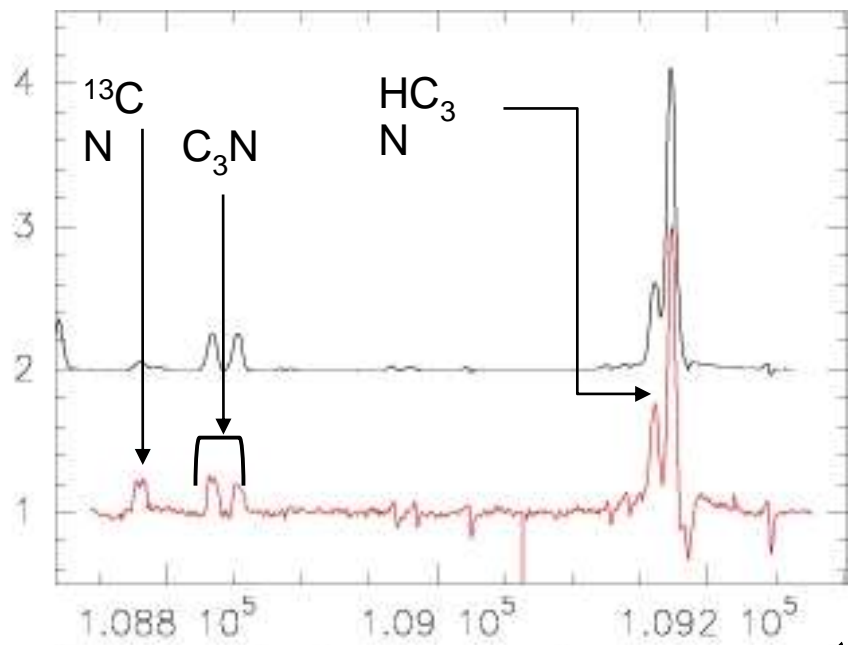
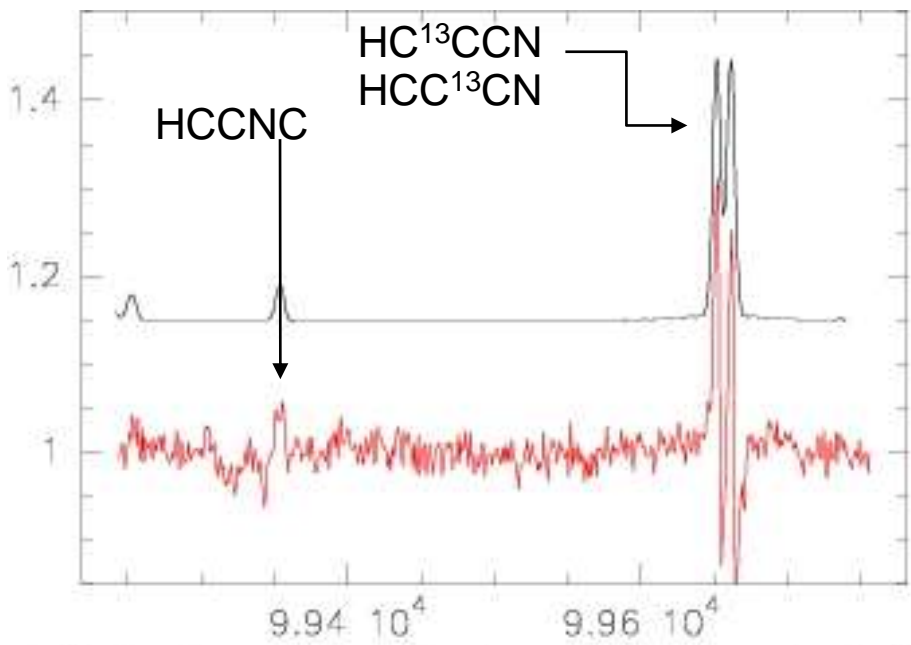
PPNe - Model

- CRL 618
 - Input
 - Temperature structure in slowly expanding envelope (SEE)
 - Number density of HC_3N
 - Output
 - Spectrum
 - Column density of HC_3N

— model
— data



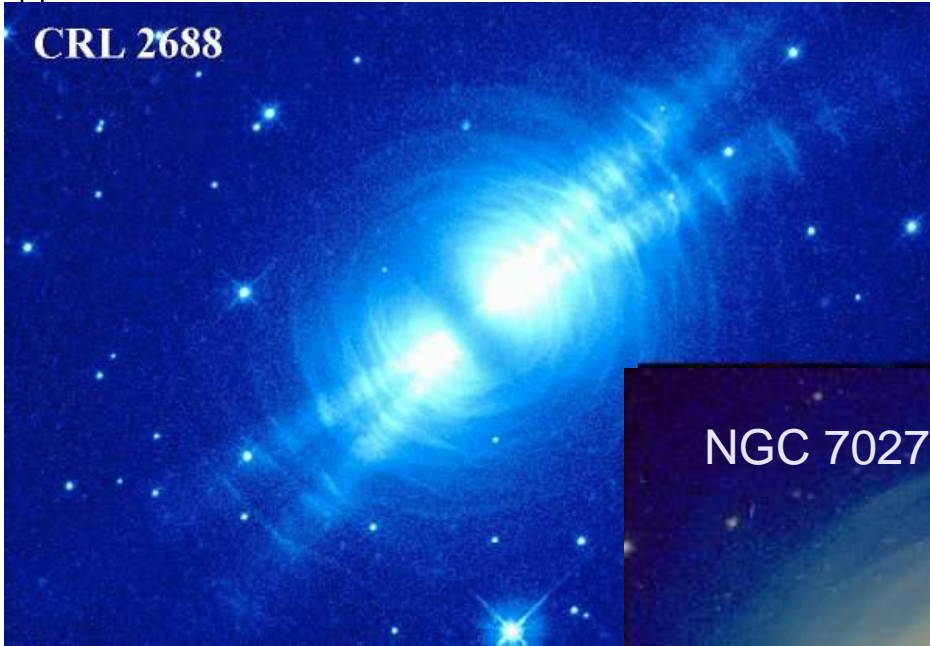
— model
— data



Protoplanetary Nebulae (PPNe)

comparing CRL 2688, CRL 618 and NGC 7027

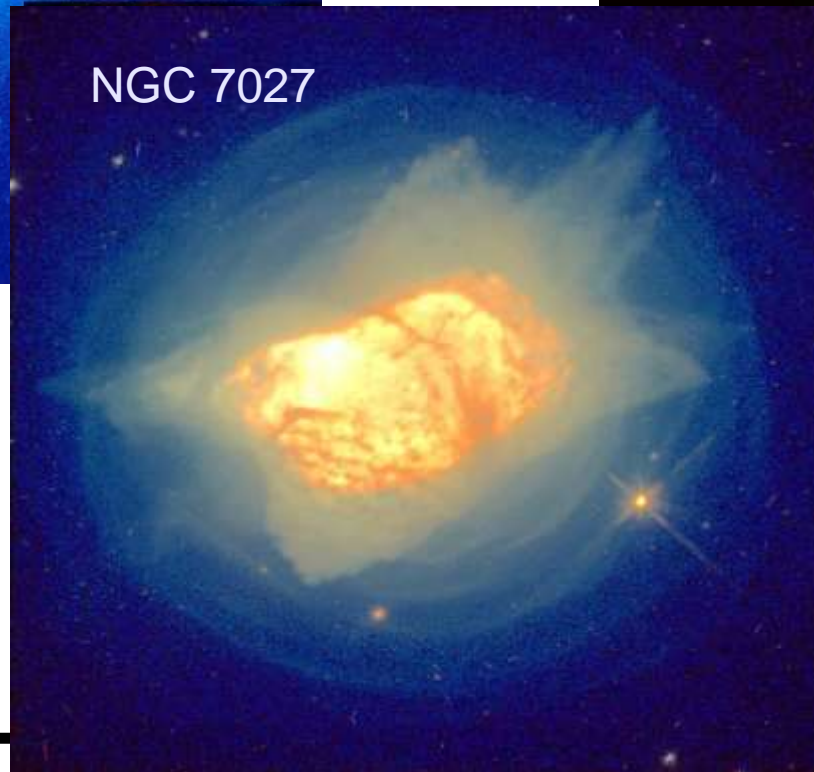
CRL 2688



CRL618



NGC 7027



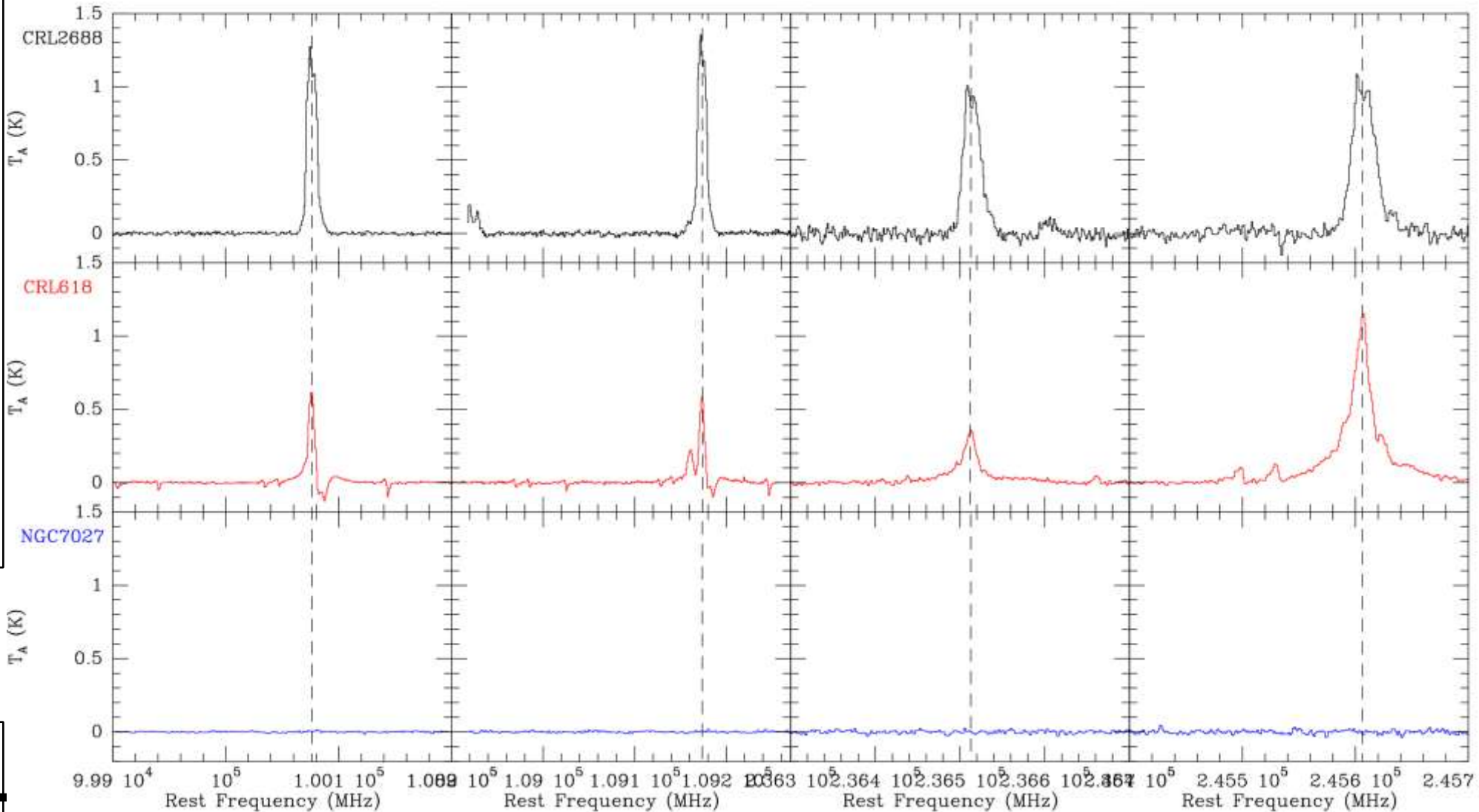
Comparing CRL 2688, CRL 618 and NGC 7027

HC₃N (11-10)

HC₃N (12-11)

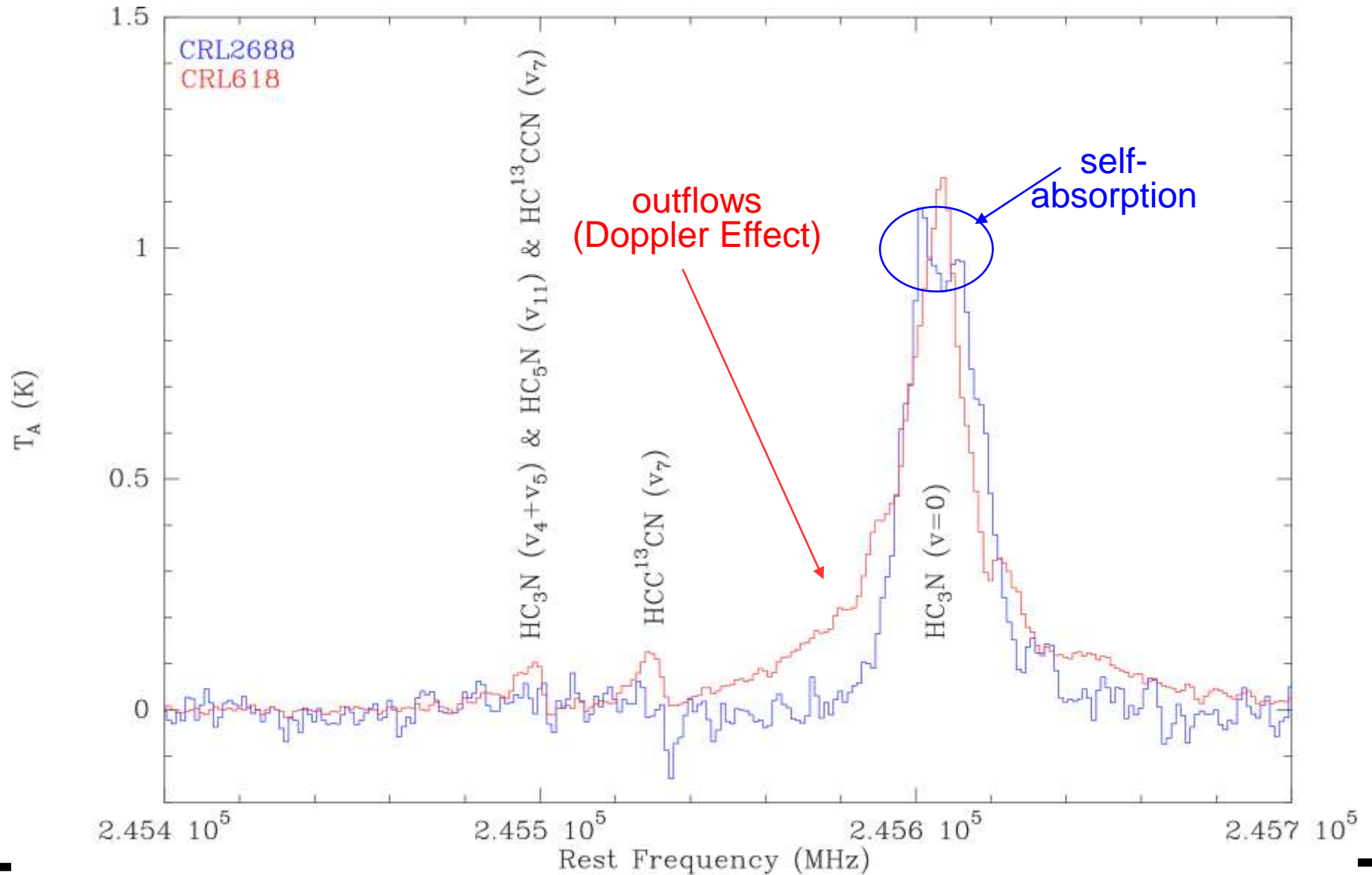
HC₃N (26-25)

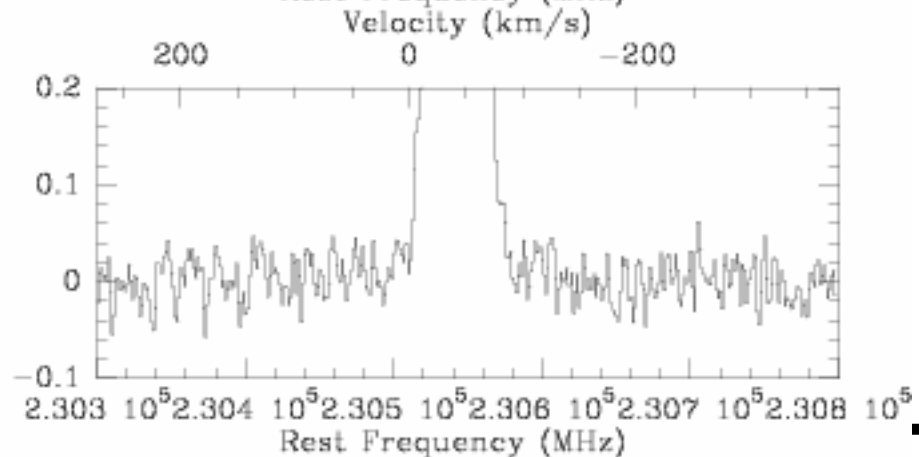
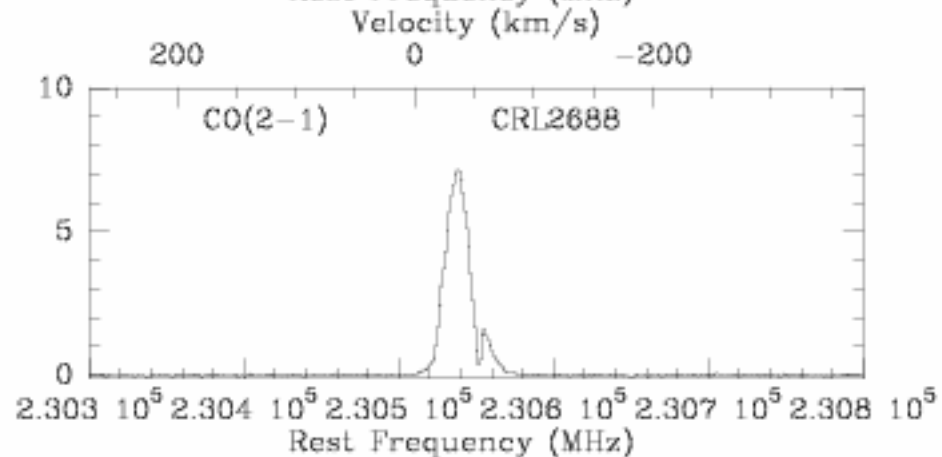
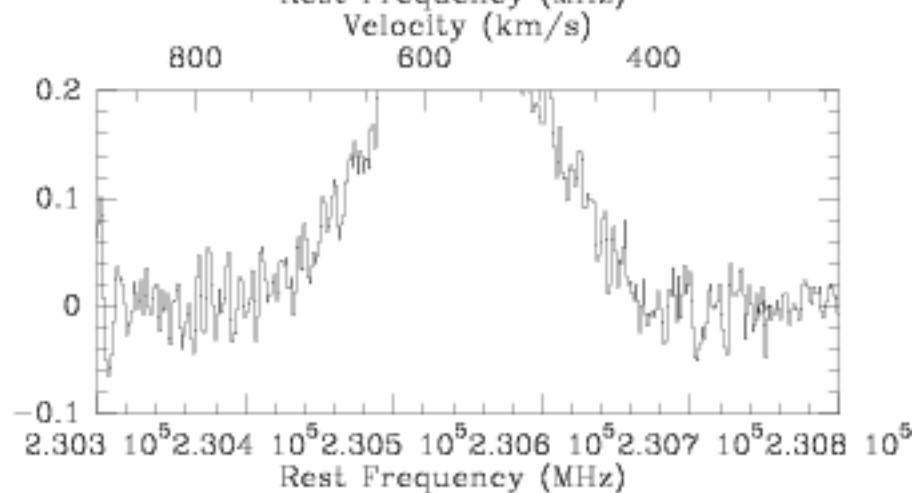
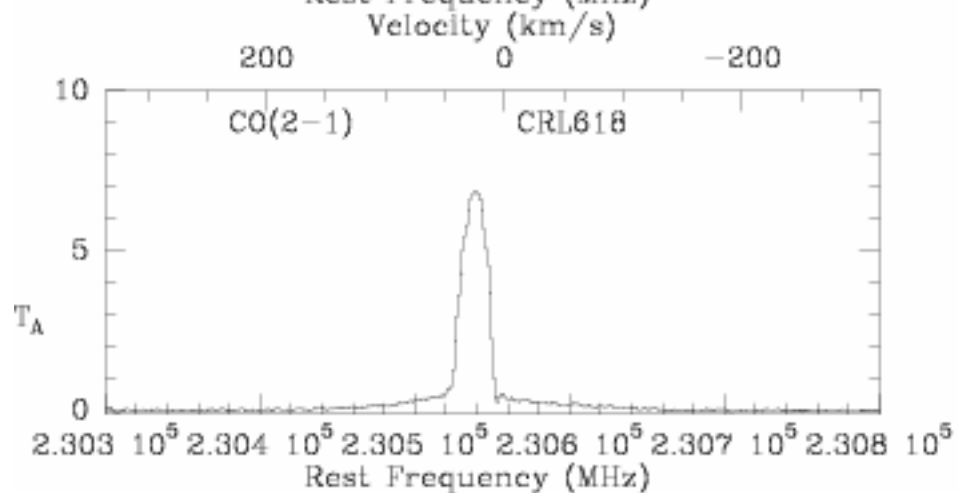
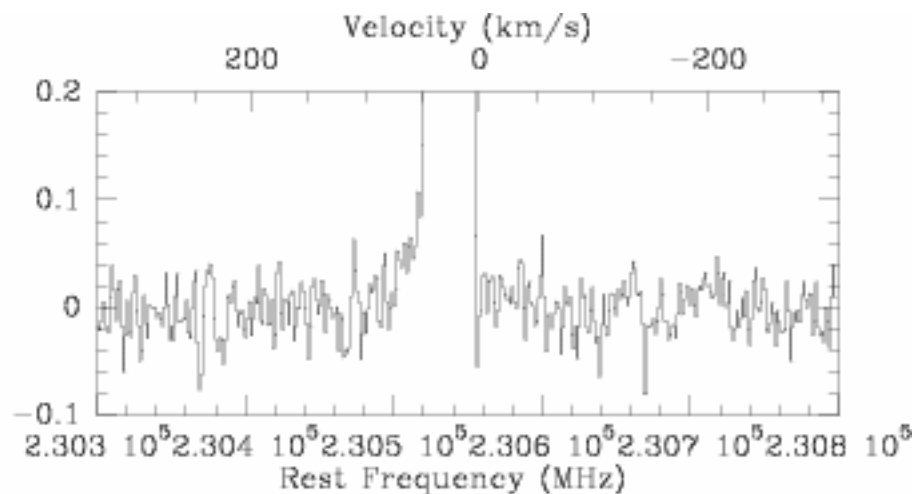
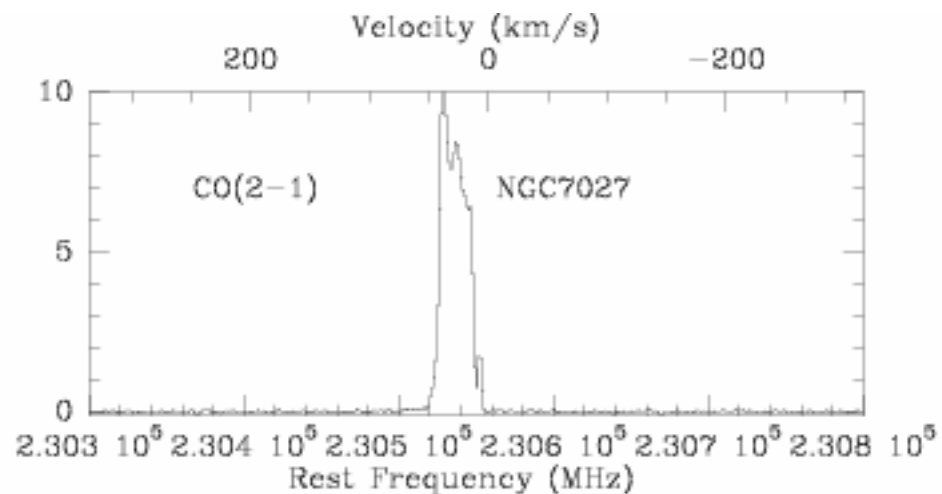
HC₃N (27-26)



Evolution & increasing Temperature

Comparing CRL 2688 & CRL 618





THANK YOU ALL



AGB stars

- outflow

- dense, molecular
- low velocities ($< 30 \text{ km s}^{-1}$)
- \dot{M} up to $10^{-5} M_{\text{sol}}/\text{yr}$

- many molecular lines to probe structure and dynamics

Sources under study

- R Aql
 - χ Cyg
- } → C-rich Miras

